# The Riesz energy of point sets on the unit sphere under weak-star convergence

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## The players: the sphere, point sets, potentials

The unit sphere is  $\mathbb{S}^d := \left\{ x \in \mathbb{R}^{d+1} \mid \|x\| = 1 
ight\}$  .

We treat sequences of point sets of the form  $(X_1, X_2, \ldots)$ , with

$$X_n := \{x_{n,1}, \ldots, x_{n,n}\} \subset \mathbb{S}^d$$
, so that  $|X_n| = n$ ,

the Riesz potential

$$U_s(r):=r^{-s}, \quad 0 < s < d,$$

and Riesz kernel

$$\Phi_s(x,y) := U_s(\|x-y\|) = \|x-y\|^{-s}.$$

For a sequence of point sets X on  $\mathbb{S}^d$ , define

$$\mathrm{E}_n(X) \Phi_s := rac{1}{n^2} \sum_{k=1}^n \sum_{\substack{j=1 \ j 
eq k}}^n \Phi_s(x_{n,k}, x_{n,j}).$$

Let  $\mu$  be the normalized Lebesgue area measure on  $\mathbb{S}^d$ . Define

$$\mathrm{I}\,\Phi_s := \int \int \Phi_s(x,y) \ d\mu(x) \ d\mu(y).$$

**Theorem 1.** For a well-separated and weak-star convergent sequence of point sets X on  $\mathbb{S}^d$ , for 0 < s < d,

$$\mathbf{E}_n(X)\Phi_s \to \mathbf{I}\Phi_s \quad as \quad n \to \infty.$$



We say our sequence X is well separated

if there is a constant  $\gamma$  such that

$$\|x-y\| > \gamma n^{-1/d}$$

for all  $x, y \in X_n$ .

A sequence of measures  $(\nu_1, \nu_2, \ldots)$ 

on a compact metric space S

converges weakly to the measure  $\nu$  if and only if

$$\int_S f(x) \ d
u_n(x) 
ightarrow \int_S f(x) \ d
u(x)$$

for all continuous f.

(Billingsley 1968, 1999)

Our point set  $X_n \subset \mathbb{S}^d$  defines a normalized counting measure via equal weight quadrature,

$$\mu_n(A) := rac{|X_n \cap A|}{|X_n|} = rac{|X_n \cap A|}{|n|}, \ \int_{\mathbb{S}^d} f(x) \ d\mu_n(x) = rac{1}{n} \sum_{k=1}^n f(x_{n,k}).$$

We say that X is weak-star convergent if  $\mu_n$  converges weakly to  $\mu$ , the normalized Lebesgue area measure on  $\mathbb{S}^d$ .

Let  $\Omega_s$  be a sequence of minimum  $\mathbf{E}_n(.)\Phi_s$  energy point sets on  $\mathbb{S}^d$ , for 0 < s < d. Then

$$\begin{split} \mathbf{E}_{n}(\Omega_{s})\Phi_{s} &= \mathbf{I}\,\Phi_{s} + O(n^{s/d-1}) \quad \text{and} \\ \mathbf{E}_{n}(\Omega_{s})\Phi_{s} &\leq \mathbf{I}\,\Phi_{s} - Cn^{s/d-1}. \end{split}$$

(Wagner 1990, 1992, Rakhmanov, Saff and Zhou 1994, Brauchart 2004)

For  $d - 1 \leq s < d$ ,  $\Omega_s$  is well separated. (Dahlberg 1978, Kuijlaars, Saff and Sun 2004)

It is well known that for 0 < s < d,  $\Omega_s$  is weak-star convergent. (Götz 2000, 2003, Götz and Saff 2000, Damelin and Grabner 2003)

## **Examples**

1. Minimum energy sequences.

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For d-1 \leqslant s' < d and 0 < s < d,\mathrm{E}_n(\Omega_{s'}) \Phi_s 	o \mathrm{I} \, \Phi_s.
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- Well-separated, diameter-bounded equal area sequences. (Alexander 1972, Stolarsky 1973, Rakhmanov, Saff and Zhou 1994, Zhou 1995, Kuijlaars and Saff 1998)
- 3. Well-separated spherical designs.(Grabner and Tichy 1993, Hesse and Leopardi (submitted))
  - For strength t, spherical cap discrepancy is  $O(t^{-1})$ .

For a measure  $\nu$  and a point x,

define the punctured measure  $\boldsymbol{\nu}[\boldsymbol{x}]$  by

$$u[x](A):=
u\left(A\setminus\{x\}
ight).$$

Then, eg.

$$\int_{\mathbb{S}^d} f(y) \ d\mu_n[x_{n,k}](y) = rac{1}{n} \sum_{\substack{j=1\ j 
eq k}}^n f(x_{n,j}).$$

For a measure u and a kernel  $\phi$  defined on a space S define

$$egin{aligned} \mathcal{I}(
u)\phi &:= \int_S \int_S \phi(x,y) \ d
u(y) \ d
u(x), \ \mathcal{E}(
u)\phi &:= \int_S \int_S \phi(x,y) \ d
u[x](y) \ d
u(x). \end{aligned}$$

For  $\mathbb{S}^d$  and the measures  $\mu$ , and  $\mu_n$  for a fixed X, define

$$egin{aligned} &\mathrm{I}_n := \mathcal{I}(\mu_n), &\mathrm{I} := \mathcal{I}(\mu), \ &\mathrm{E}_n := \mathcal{E}(\mu_n), &\mathrm{E} := \mathcal{E}(\mu). \end{aligned}$$

These definitions agree with those used in Theorem 1.

Fix 0 < s < d.

Fix a weak-star convergent and well-separated sequence X,

with separation constant  $\gamma$ .

Define  $\Delta_m := \gamma m^{-1/d}$ .

Now define the cutoff potential  $U_m$  by

$$U_m(r) := egin{cases} \Delta_m^{-s} & (0 < r \leqslant \Delta_m) \ U(r) = r^{-s} & (r > \Delta_m). \end{cases}$$

By the triangle inequality, for any  $m, n \in \mathbb{N}$ ,

$$|\mathrm{E}_n\,\Phi-\mathrm{I}\,\Phi|\leqslant a_n+b_{m,n}+c_{m,n}+d_{m,n}+e_m,$$

where

$$egin{aligned} &a_n := |\mathrm{E}_n\,\Phi - \mathrm{E}_n\,\Phi_n|,\ &b_{m,n} := |\mathrm{E}_n\,\Phi_n - \mathrm{E}_n\,\Phi_m|,\ &c_{m,n} := |\mathrm{E}_n\,\Phi_m - \mathrm{I}_n\,\Phi_m|,\ &d_{m,n} := |\mathrm{I}_n\,\Phi_m - \mathrm{I}\,\Phi_m|,\ &e_m := |\mathrm{I}\,\Phi_m - \mathrm{I}\,\Phi_m|. \end{aligned}$$

Convergence of 
$$a_n = |\mathrm{E}_n(\Phi - \Phi_n)|$$

Since 
$$(U - U_n)(r) = 0$$
 for  $r \ge \Delta_n$ ,

the separation condition guarantees that

 $a_n = 0$  for any n.

Convergence of 
$$b_{m,n} = |\mathrm{E}_n(\Phi_n - \Phi_m)|$$

For  $m \ge n$ ,  $b_{m,n} = 0$  by the same argument as for  $a_n$ .

Define  $\Theta_n := 2 \sin^{-1}(\Delta_n/2)$ .

For m < n, for any k, separation gives a bound on

$$\mu_nig(S(x_{n,k}, heta)\setminus S(x_{n,k},\Theta_n)ig)$$

in terms of the packing of small spherical caps  $S(y, \Theta_n/2)$  within the larger cap  $S(x_{n,k}, \theta + \Theta_n/2)$ .

We use this bound with a Riemann-Stieltjes integral to show that  $b_{m,n} < \frac{\epsilon}{4}$  if both m and  $n \ge M_b$ , where  $M_b$  is defined using  $\epsilon, d, s$  and  $\gamma$ .

Convergence of 
$$c_{m,n} = |(\mathrm{E}_n - \mathrm{I}_n)\Phi_m|$$

$$egin{aligned} m{c}_{m,n} &= rac{1}{n^2} \sum_{k=1}^n \Phi_m(x_{n,k},x_{n,k}) \ &= rac{U_m(0)}{n} = rac{\Delta_m^{-s}}{n} = \gamma^{-s} m^{s/d} n^{-1} \ &< rac{\epsilon}{4} & ext{if} \quad n \geqslant N_c(m), \end{aligned}$$

#### where

$$N_c(m) := 4\gamma^{-s}m^{s/d}\epsilon^{-1} + 1.$$

For any m, the weak-star convergence of X ensures that

$$d_{m,n} o 0$$
 as  $n o \infty,$ 

since weak convergence of  $\mu_n \to \mu$ implies weak convergence of  $\mu_n \times \mu_n \to \mu \times \mu$ .

(Billingsley 1968, 1999)

So we have  $d_{m,n} < \frac{\epsilon}{4}$  if  $n \ge N_d(m)$ , where  $N_d(m)$  depends on  $\epsilon, m$  and X.

Convergence of 
$$e_m = |I(\Phi_m - \Phi)|$$

We have

$$e_m = rac{\omega_d}{\omega_{d+1}} \int_0^{\Theta_m} (r^{-s} - \Delta_m^{-s}) (\sin \theta)^{d-1} d\theta$$

where 
$$\omega_d := \sigma(\mathbb{S}^{d-1})$$
 and  $r = 2\sin(\theta/2)$ .

It can be shown that  $e_m < \frac{\epsilon}{4}$  if  $m \ge M_e$ , where  $M_e$  is defined using  $\epsilon, d, s$  and  $\gamma$ . So we have  $|\mathbf{E}_n \Phi - \mathbf{I} \Phi| < \epsilon$  if

$$n \geqslant \max\left(M, N_{oldsymbol{c}}(M), N_{oldsymbol{d}}(M)
ight),$$

where

$$M := \max(M_b, M_e).$$

In other words,  $|\mathbf{E}_n \Phi - \mathbf{I} \Phi| \to 0$  as  $n \to \infty$ .