Ramanujan and Euler's Constant

Richard P. Brent MSI, ANU

8 July 2010

In memory of Ed McMillan 1907 – 1991

Presented at the CARMA Workshop on *Exploratory Experimentation and Computation in Number Theory*, Newcastle, Australia, 7–9 July 2010.

Copyright © 2010, R. P. Brent

Motivation

Ramanujan gave many beautiful formulas for π and $1/\pi$. See, for example, J. M. Borwein and P. B. Borwein, *Pi and the AGM*, John Wiley and Sons, New York, 1987; also (same authors) "Ramanujan and Pi", *Scientific American*, February 1988, 66–73.

Euler's constant

$$\gamma = -\Gamma'(1) \simeq 0.577$$

is more mysterious than π . For example, unlike π , we do not know any quadratically convergent iteration for γ . We do not know if γ is transcendental. We do not even know if γ is irrational, though this seems likely. All we know is that if $\gamma = p/q$ is rational, then *q* is large. This follows from a computation of the regular continued fraction expansion for γ .

(日) (圖) (E) (E) (E)

Apéry proved $\zeta(3)$ irrational using the series

$$\zeta(3) = \frac{5}{2} \sum_{k=1}^{\infty} \frac{(-1)^{k-1} k! k!}{(2k)! k^3}$$

and, in Chapter 9 of his Notebooks, Ramanujan gives several similar series, some involving $\zeta(3)$.

< □→ < □→ < □→ - □

Ramanujan rediscovered Euler's formula

$$\zeta(\mathbf{3}) = \sum_{k=1}^{\infty} \frac{H_k}{(k+1)^2},$$

where

$$H_k = \sum_{j=1}^k \frac{1}{j}$$

is a Harmonic number. Harmonic numbers also occur in formulas involving γ (examples later).

Thus, it is natural to look in the work of Ramanujan for formulas involving γ , in the hope that some of these might be useful for computing accurate approximations to γ , or even for proving that γ is irrational.

<回> < 回> < 回> < 回> = □

Ramanujan's Papers and Notebooks

Ramanujan published one paper specifically on γ : "A series for Euler's constant γ ", *Messenger of Mathematics* 46 (1917), 73–80. In this paper he generalizes an interesting series which was first discovered by Glaisher:

$$\gamma = 1 - \sum_{k=1}^{\infty} \frac{\zeta(2k+1)}{(k+1)(2k+1)}.$$

This family of series all involve the Riemann zeta function or related functions, so they are not convenient for computational purposes.

Much of Ramanujan's work was not published during his lifetime, but was summarized in his Notebooks. Edited editions have been published by Berndt [1]. In the following, page numbers refer to Berndt's edition (Part I for Chapters 1–9, Part II for Chapters 10–15).

< □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □

γ in Ramanujan's Notebooks

Scanning Berndt, we find many occurrences of γ . Some involve the logarithmic derivative $\psi(x)$ of the gamma function, or the sum

$$H_x = \sum_{k=1}^n 1/k,$$

which we can interpret as $\psi(x + 1) + \gamma$ if x is not necessarily a positive integer (Ch. 8, pg. 181). There are also applications of the result

$$H_n = \ln n + \gamma + O(1/n)$$

as $n \to \infty$.

Other interesting formulas involving γ occur in Chapters 14–15, e.g. Ch. 15, Entry 1, examples (i–ii), pp. 303–304.

(日) (圖) (E) (E) (E)

Chapter 4, Entry 9

We shall concentrate on Chapter 4, Entry 9, Corollaries 1–2 (pg. 98), because these are potentially useful for computing γ . Corollary 1 is

$$\sum_{k=1}^{\infty} \frac{(-1)^{k-1} x^k}{k! k} \sim \ln x + \gamma \tag{1}$$

as $x \to \infty$. In fact, Euler showed that

$$\sum_{k=1}^{\infty} \frac{(-1)^{k-1} x^k}{k! k} - \ln x - \gamma = \int_x^{\infty} \frac{e^{-t}}{t} dt = O\left(\frac{e^{-x}}{x}\right)$$

and this has been used by Sweeney and others to compute Euler's constant (one has to be careful because of cancellation in the series). In Ch. 12, Entry 44(ii), Ramanujan states Euler's result that the error is between $e^{-x}/(1+x)$ and e^{-x}/x .

< □→ < □→ < □→ - □

A Generalization

Ramanujan's Corollary 2, Entry 9, Chapter 4 (page 98) is that, for positive integer *n*,

$$\sum_{k=1}^{\infty} \frac{(-1)^{k-1}}{nk} \left(\frac{x^k}{k!}\right)^n \sim \ln x + \gamma$$
 (2)

so (1) is just the case n = 1.

Berndt shows that (2) is false for $n \ge 3$. In fact, the function defined by the left side of (2) changes sign infinitely often, and grows exponentially large as $x \to \infty$. However, Berndt leaves the case n = 2 open.

< □→ < □→ < □→ = □

We shall sketch a proof that (2) is true in the case n = 2. In fact, we shall obtain an exact expression for the error in (2) as an integral involving the Bessel function $J_0(x)$, and deduce an asymptotic expansion.

The exact expression for n = 2 is a special case of a formula given on page 48 of Y. L. Luke, *Integrals of Bessel Functions*, 1962. However, Luke does not comment on the connection with Ramanujan.

< □→ < □→ < □→ □ □

Avoiding Cancellation

In Chapter 3, Entry 2, Cor. 2, page 46, Ramanujan states that the sum

$$\sum_{k=1}^{\infty} \frac{(-1)^{k-1} x^k}{k! k}$$

occurring in (1) can be written as

$$e^{-x}\sum_{k=0}^{\infty}H_k\frac{x^k}{k!}.$$

This is easy to prove (Berndt, page 47). Thus (1) gives

$$\sum_{k=0}^{\infty} H_k \frac{x^k}{k!} \Big/ \sum_{k=0}^{\infty} \frac{x^k}{k!} \sim \ln x + \gamma.$$
(3)

This is more convenient than (1) for computation, because there is no cancellation in the series when x > 0. Later we indicate how Ramanujan might have generalized (3) in much the same way that he attempted to generalize (1),

Ramanujan's Corollary for n = 2

The following result from [3] shows that (2) is valid for n = 2. Recall that

$$J_0(x) = \sum_{k=0}^{\infty} \frac{(-1)^k (x/2)^{2k}}{k! k!}$$

is a Bessel function of the first kind and order zero.

Theorem

Let

$$e(x) = \sum_{k=1}^{\infty} \frac{(-1)^{k-1}}{2k} \left(\frac{x^k}{k!}\right)^2 - \ln x - \gamma.$$

Then, for real positive x,

$$e(x) = \int_{2x}^{\infty} \frac{J_0(t)}{t} dt.$$

< □→ < □→ < □→ □ □

Sketch of Proof

Proceed as on pg. 99 of Berndt, and use the fact that

$$\int_0^\infty \left(\frac{e^{-t} - J_0(2t)}{t}\right) dt = 0.$$
(4)

A slightly more general result than (4) is given in equation 6.622.1 of Gradshteyn and Ryzhik, and is attributed to Nielsen. An independent proof is given in [3].

Corollary

Let e(x) be as in Theorem 1. Then, for large positive x, e(x) has an asymptotic expansion

$$e(x) = \frac{1}{2\pi^{1/2}x^{3/2}} \left(\cos\left(2x + \frac{\pi}{4}\right) + \frac{13\sin\left(2x + \frac{\pi}{4}\right)}{16x} + O\left(\frac{1}{x^2}\right) \right).$$

▲□ → ▲ □ → ▲ □ → □ □

Comparison of n = 1 and n = 2

We see that, for computational purposes, it is much better to take n = 1 than n = 2 in (2), because the error for n = 1 is $O(e^{-x}/x)$, but for n = 2 it is $\Omega_{\pm}(x^{-3/2})$.

< 国 > < 国 > < 国 > -

크

A Different Generalization

We obtained (2) from (1) by replacing $x^k/k!$ by $(x^k/k!)^n/n$. A similar generalization of (3) is

$$\sum_{k=0}^{\infty} H_k \left(\frac{x^k}{k!}\right)^n \Big/ \sum_{k=0}^{\infty} \left(\frac{x^k}{k!}\right)^n \sim \ln x + \gamma$$
 (5)

as $x \to \infty$. (3) is just the case n = 1.

It is easy to show that (5) is valid for all positive integer *n*. An essential difference between (2) and (5) is that there is a large amount of cancellation between terms on the left side of (2), but there is no cancellation in the numerator and denominator on the left side of (5). The function $(x^k/k!)^n$ acts as a smoothing kernel with a peak at $k \simeq x$. Since

$$H_k = \ln k + \gamma + O(1/k),$$

the result (5) is not surprising. What may be surprising is the speed of convergence.

Speed of Convergence

Brent and McMillan [2] show that

$$\sum_{k=0}^{\infty} H_k \left(\frac{x^k}{k!}\right)^n / \sum_{k=0}^{\infty} \left(\frac{x^k}{k!}\right)^n = \ln x + \gamma + O(e^{-c_n x}) \qquad (6)$$

as $x \to \infty$, where

$$c_n = egin{cases} 1, & ext{if } n = 1; \ 2n \sin^2(\pi/n), & ext{if } n \geq 2. \end{cases}$$

In the case n = 2, (6) has error $O(e^{-4x})$. Brent and McMillan used this case with $x \simeq 17,400$ to compute γ to high precision. They deduced that, if $\gamma = p/q$ is rational, then $q > 10^{15000}$. From Corollary 1, the same value of x in (2) would give less than 8-decimal place accuracy.

Another view

If you are looking for a good way to compute Euler's constant γ , you might scan Abramowitz and Stegun (or the online *Digital Library of Mathematical Functions*) looking for formulas in which γ occurs.

For example, in the chapter on Bessel functions, we find (9.6.13):

$$K_0(2x) = -(\ln(x) + \gamma)I_0(2x) + \frac{x^2}{(1!)^2} + (1 + \frac{1}{2})\frac{x^4}{(2!)^2} + \cdots$$

(where I replaced z by 2x). Here $I_0(z)$ and $K_0(z)$ are modified Bessel functions (sometimes called Bessel functions of imaginary argument because we obtain them by $z \mapsto iz$ in the usual Bessel functions $J_0(z)$ etc).

◆□▶ ◆□▶ ◆三▶ ◆三▶ ◆□▶

Another view

If you are looking for a good way to compute Euler's constant γ , you might scan Abramowitz and Stegun (or the online *Digital Library of Mathematical Functions*) looking for formulas in which γ occurs.

For example, in the chapter on Bessel functions, we find (9.6.13):

$$\mathcal{K}_{0}(2x) = -(\ln(x) + \gamma)I_{0}(2x) + \frac{x^{2}}{(1!)^{2}} + (1 + \frac{1}{2})\frac{x^{4}}{(2!)^{2}} + \cdots$$

(where I replaced z by 2x). Here $I_0(z)$ and $K_0(z)$ are modified Bessel functions (sometimes called Bessel functions of imaginary argument because we obtain them by $z \mapsto iz$ in the usual Bessel functions $J_0(z)$ etc).

Formula (9.6.13) might be useful for evaluating γ if we had an independent way of evaluating $K_0(2x)$ and $I_0(2x)$.

「日本日本日本」 日

Differential equation (ODE)

 $I_0(z)$ and $K_0(z)$ are independent solutions of the *modified* Bessel equation

$$zw''+w'-zw=0.$$

This is the special case $\nu = 0$ of

$$z^2w'' + zw' - (z^2 + \nu^2)w = 0.$$

<回><モン<

크

Abramowitz and Stegun (9.6.12) gives the nice series

$$I_0(2x) = \sum_{k=0}^{\infty} \frac{x^{2k}}{(k!)^2},$$

so there is no difficulty in computing $I_0(2x)$. (As usual, we replaced *z* by 2*x*.)

日本・日本・日本・

2

Asymptotic series for I_0 and K_0

In the same chapter of Abramowitz and Stegun, we find the asymptotic expansions:

$$\begin{split} & l_0(z) \sim \frac{e^z}{\sqrt{2\pi z}} \left(1 + \frac{1^2}{1!(8z)} + \frac{1^2 \cdot 3^2}{2!(8z)^2} + \frac{1^2 \cdot 3^2 \cdot 5^2}{3!(8z)^3} + \cdots \right), \\ & \mathcal{K}_0(z) \sim e^{-z} \sqrt{\frac{\pi}{2z}} \left(1 - \frac{1^2}{1!(8z)} + \frac{1^2 \cdot 3^2}{2!(8z)^2} - \frac{1^2 \cdot 3^2 \cdot 5^2}{3!(8z)^3} + \cdots \right). \end{split}$$

These expansions give a way of computing $I_0(z)$ and $K_0(z)$ accurately if z is sufficiently large (z is always real and positive in our applications).

▲御 ▶ ▲ 臣 ▶ ▲ 臣 ▶ 二 臣

Asymptotic series for I_0 and K_0

In the same chapter of Abramowitz and Stegun, we find the asymptotic expansions:

$$\begin{split} & l_0(z) \sim \frac{e^z}{\sqrt{2\pi z}} \left(1 + \frac{1^2}{1!(8z)} + \frac{1^2 \cdot 3^2}{2!(8z)^2} + \frac{1^2 \cdot 3^2 \cdot 5^2}{3!(8z)^3} + \cdots \right), \\ & \mathcal{K}_0(z) \sim e^{-z} \sqrt{\frac{\pi}{2z}} \left(1 - \frac{1^2}{1!(8z)} + \frac{1^2 \cdot 3^2}{2!(8z)^2} - \frac{1^2 \cdot 3^2 \cdot 5^2}{3!(8z)^3} + \cdots \right). \end{split}$$

These expansions give a way of computing $I_0(z)$ and $K_0(z)$ accurately if *z* is sufficiently large (*z* is always real and positive in our applications).

The leading terms show that $K_0(z)/I_0(z) = O(e^{-2z})$ is exponentially small if *z* is large and positive.

(日)

Asymptotic series for $I_0 K_0$

While on the subject of asymptotic expansions, note that if we multiply the asymptotic expansions for $I_0(z)$ and $K_0(z)$, then half the terms vanish, and we obtain (at least formally)

$$I_0(z)K_0(z) \sim \frac{1}{2z} \left(1 + \frac{1^3}{1!(8z^2)} + \frac{1^3 \cdot 3^3}{2!(8z^2)^2} + \frac{1^3 \cdot 3^3 \cdot 5^3}{3!(8z^2)^3} + \cdots \right)$$

This turns out to be in Abramowitz and Stegun (9.7.5), though no error bound or reference is given there.

伺下 イヨト イヨト

Asymptotic series for $I_0 K_0$

While on the subject of asymptotic expansions, note that if we multiply the asymptotic expansions for $I_0(z)$ and $K_0(z)$, then half the terms vanish, and we obtain (at least formally)

$$I_0(z)K_0(z) \sim \frac{1}{2z} \left(1 + \frac{1^3}{1!(8z^2)} + \frac{1^3 \cdot 3^3}{2!(8z^2)^2} + \frac{1^3 \cdot 3^3 \cdot 5^3}{3!(8z^2)^3} + \cdots \right)$$

This turns out to be in Abramowitz and Stegun (9.7.5), though no error bound or reference is given there.

To prove the formula for the general term, we could use the Wilf-Zeilberger (WZ) method. Easier is to use the ODE

$$z^3 f''' + z(1 - 4z^2)f' - f = 0$$

satisfied by $f(z) = zI_0(z)K_0(z)$. It is straightforward to deduce a recurrence relation for the coefficients in the asymptotic expansion from this ODE.

▲□ → ▲ □ → ▲ □ → □ □

Error bounds for the asymptotic expansions

Suppose the asymptotic expansions for I_0 , K_0 or I_0K_0 are written as

$$F(z) \sim a_0(z) + a_1(z) + a_2(z) + \cdots$$

(where the $a_j(z)$ are not identically zero), and the error $E_n(z)$ is defined by

$$F(z) = a_0(z) + a_1(z) + \cdots + a_{n-1}(z) + E_n(z).$$

Then, provided z is real, $z \ge 1$, and n > 0, we can show that

$$|E_n(z)| = O(\sqrt{n}|a_n(z)|),$$

and even give an explicit constant in the "O" result (e.g. 4). In the case of K_0 , the errors alternate in sign and there is a sharper bound

$$|E_n(z)| < |a_n(z)|.$$

◆□▶ ◆□▶ ◆三▶ ◆三▶ ◆□▶

Proofs of Error Bounds

The proofs for K_0 and I_0 are discussed in Olver's book (Chapter 7, especially Ex. 13.2).

▲□ → ▲ □ → ▲ □ → …

2

The proofs for K_0 and I_0 are discussed in Olver's book (Chapter 7, especially Ex. 13.2).

The proof for $I_0 K_0$ does not seem to have been published, and was stated as a conjecture in Brent and McMillan (1980). It is possible to deduce it from bounds for the K_0 and I_0 expansions.

伺 ト イ ヨ ト イ ヨ ト

Deducing γ

Rearranging (9.6.13) and using the power series for I_0 gives

$$\gamma + \ln(x) = \sum_{k=0}^{\infty} H_k \left(\frac{x^k}{k!}\right)^2 / \sum_{k=0}^{\infty} \left(\frac{x^k}{k!}\right)^2 - \frac{K_0(2x)}{I_0(2x)},$$

but the last term is $O(e^{-4x})$ so can be neglected if *x* is large. This is essentially Algorithm B1 of Brent & McMillan.

< 国 > < 国 > < 国 > -

크

Deducing γ

Rearranging (9.6.13) and using the power series for I_0 gives

$$\gamma + \ln(x) = \sum_{k=0}^{\infty} H_k \left(\frac{x^k}{k!}\right)^2 / \sum_{k=0}^{\infty} \left(\frac{x^k}{k!}\right)^2 - \frac{K_0(2x)}{I_0(2x)},$$

but the last term is $O(e^{-4x})$ so can be neglected if *x* is large. This is essentially Algorithm B1 of Brent & McMillan.

We have just recovered (6) in the case n = 2, with an explicit error term $K_0(2x)/I_0(2x)$.

▲□ ▶ ▲ 臣 ▶ ▲ 臣 ▶ ● ① ● ○ ● ●

Approximating the error term

To get a more accurate algorithm (with the same *x*) we can try to approximate the error term $K_0(2x)/I_0(2x)$. Since $I_0(2x)$ has already been computed (denominator of the main term), we only need to approximate $K_0(2x)$. This can be done with relative error $O(e^{-4x})$ by taking z = 2x and $\lceil 4x \rceil$ terms in the asymptotic expansion for $K_0(z)$.

< □→ < □→ < □→ □ □

Approximating the error term

To get a more accurate algorithm (with the same *x*) we can try to approximate the error term $K_0(2x)/I_0(2x)$. Since $I_0(2x)$ has already been computed (denominator of the main term), we only need to approximate $K_0(2x)$. This can be done with relative error $O(e^{-4x})$ by taking z = 2x and $\lceil 4x \rceil$ terms in the asymptotic expansion for $K_0(z)$.

A faster way is to take $\lceil 2x \rceil$ terms in the asymptotic expansion for $I_0(2x)K_0(2x)$, and divide the result by $I_0(2x)^2$.

→ □ → モ → モ → モ → ○ ○ ○

Approximating the error term

To get a more accurate algorithm (with the same *x*) we can try to approximate the error term $K_0(2x)/I_0(2x)$. Since $I_0(2x)$ has already been computed (denominator of the main term), we only need to approximate $K_0(2x)$. This can be done with relative error $O(e^{-4x})$ by taking z = 2x and $\lceil 4x \rceil$ terms in the asymptotic expansion for $K_0(z)$.

A faster way is to take $\lceil 2x \rceil$ terms in the asymptotic expansion for $I_0(2x)K_0(2x)$, and divide the result by $I_0(2x)^2$.

In this way we get an algorithm for γ with error $O(e^{-8x})$ (Algorithm B3 of Brent & McMillan).

◆□▶ ◆□▶ ◆三▶ ◆三▶ ◆□▶

When summing series of rational numbers, the complexity can be reduced by the technique of *binary splitting*.

イロト イヨト イヨト イヨト 三日

When summing series of rational numbers, the complexity can be reduced by the technique of *binary splitting*.

Divide the sum into two, take out common factors, and recursively sum each half. This way we work with "small" rational numbers most of the time. Use "fast" integer multiplication for the rational arithmetic. If the numerators and denominators of the rationals grow too large, they can be "pruned".

<回>< E> < E> < E> = E

When summing series of rational numbers, the complexity can be reduced by the technique of *binary splitting*.

Divide the sum into two, take out common factors, and recursively sum each half. This way we work with "small" rational numbers most of the time. Use "fast" integer multiplication for the rational arithmetic. If the numerators and denominators of the rationals grow too large, they can be "pruned".

It is easy to use this technique to sum the power series for $\exp(z)$ or for $l_0(z)$ when z is rational. The complexity for d digits is reduced from $O(d^2)$ to $O(d(\log d)^c)$ for some (small) constant c.

< □→ < □→ < □→ □ □

When summing series of rational numbers, the complexity can be reduced by the technique of *binary splitting*.

Divide the sum into two, take out common factors, and recursively sum each half. This way we work with "small" rational numbers most of the time. Use "fast" integer multiplication for the rational arithmetic. If the numerators and denominators of the rationals grow too large, they can be "pruned".

It is easy to use this technique to sum the power series for $\exp(z)$ or for $l_0(z)$ when z is rational. The complexity for d digits is reduced from $O(d^2)$ to $O(d(\log d)^c)$ for some (small) constant c.

It is trickier to implement binary splitting for the series involving Harmonic numbers H_k , but it can be done.

▲母 ▶ ▲ 臣 ▶ ▲ 臣 ▶ ▲ 臣 ◆ の Q @

Binary Splitting – 1D case

The recursive procedure

$$S_1(f, j, \ell) = \begin{cases} 0 \text{ if } \ell \leq 0, \\ f_j \text{ if } \ell = 1, \\ S_1(f, j, \lfloor \ell/2 \rfloor) + S_1(f, j + \lfloor \ell/2 \rfloor, \lceil \ell/2 \rceil) \text{ otherwise} \end{cases}$$

returns the sum

$$\sum_{0\leq k<\ell}f_{j+k}$$

It is easy to modify S_1 to compute the polynomial

$$\sum_{0 \le k < \ell} f_{j+k} x^{j}$$

▲□ → ▲ □ → ▲ □ → □

크

Binary Splitting – 1D case

The recursive procedure

$$S_1(f, j, \ell) = \begin{cases} 0 \text{ if } \ell \leq 0, \\ f_j \text{ if } \ell = 1, \\ S_1(f, j, \lfloor \ell/2 \rfloor) + S_1(f, j + \lfloor \ell/2 \rfloor, \lceil \ell/2 \rceil) \text{ otherwise} \end{cases}$$

returns the sum

$$\sum_{0\leq k<\ell}f_{j+k}$$

It is easy to modify S_1 to compute the polynomial

$$\sum_{0 \le k < \ell} f_{j+k} x^j$$

(details left as an exercise).

Binary Splitting – 2D case

Similarly, to compute the sum

$$\sum_{0 \leq p+q < \ell} f_{j+p} g_{k+q}$$

we can use the recursive procedure

(

$$S_2(j,k,\ell) = \begin{cases} 0 \text{ if } \ell \leq 0, \\ f_j g_k \text{ if } \ell = 1, \\ S_2(j + \lfloor \ell/2 \rfloor, k, \lceil \ell/2 \rceil) + S_2(j,k + \lceil \ell/2 \rceil, \lfloor \ell/2 \rfloor) + \\ S_1(f,j, \lfloor \ell/2 \rfloor) S_1(g,k, \lceil \ell/2 \rceil) \text{ otherwise.} \end{cases}$$

★ E ► < E ►</p>

크

Binary Splitting – 2D case

Similarly, to compute the sum

$$\sum_{0 \leq p+q < \ell} f_{j+p} g_{k+q}$$

we can use the recursive procedure

(

$$S_2(j,k,\ell) = \begin{cases} 0 \text{ if } \ell \leq 0, \\ f_j g_k \text{ if } \ell = 1, \\ S_2(j + \lfloor \ell/2 \rfloor, k, \lceil \ell/2 \rceil) + S_2(j,k + \lceil \ell/2 \rceil, \lfloor \ell/2 \rfloor) + \\ S_1(f,j, \lfloor \ell/2 \rfloor) S_1(g,k, \lceil \ell/2 \rceil) \text{ otherwise.} \end{cases}$$

This is essentially the "short product" algorithm of Mulders (2000).

伺 ト イ ヨ ト イ ヨ ト -

Binary Splitting cont.

We can use the recursive procedure S_2 to compute sums such as

$$\sum_{0 < k < n} H_k b_k = \sum_{0 < j \le k < n} \frac{b_k}{j}$$

◆□→ < □→ < □→ < □</p>

Binary Splitting cont.

We can use the recursive procedure S_2 to compute sums such as

$$\sum_{0 < k < n} H_k b_k = \sum_{0 < j \le k < n} \frac{b_k}{j}$$

(details left as an exercise).

B > 4 B > 1

2

Alexander Yee seems to hold the world record for the computation of γ (though records do not last long, so this may soon be obsolete). In March 2009 he computed the first 29844 489 545 $\approx 2^{36}/\ln(10)$ decimal digits.

▲御▶ ▲ 理▶ ▲ 理▶ ― 理

Alexander Yee seems to hold the world record for the computation of γ (though records do not last long, so this may soon be obsolete). In March 2009 he computed the first 29 844 489 545 $\approx 2^{36}/\ln(10)$ decimal digits.

The computation was performed using Algorithm B3 with binary splitting and $x = 2^{33}$ (taking 205 hours on a dual-processor workstation "Nagisa").

▲掃▶▲≣▶▲≣▶ ≣ のQ@

Alexander Yee seems to hold the world record for the computation of γ (though records do not last long, so this may soon be obsolete). In March 2009 he computed the first 29 844 489 545 $\approx 2^{36}/\ln(10)$ decimal digits.

The computation was performed using Algorithm B3 with binary splitting and $x = 2^{33}$ (taking 205 hours on a dual-processor workstation "Nagisa").

Verification used Algorithm B1 with $x = 2^{34}$ (taking 269 hours).

|▲□ ▶ ▲ 臣 ▶ ▲ 臣 ▶ ○ 臣 ○ � � �

Alexander Yee seems to hold the world record for the computation of γ (though records do not last long, so this may soon be obsolete). In March 2009 he computed the first 29 844 489 545 $\approx 2^{36}/\ln(10)$ decimal digits.

The computation was performed using Algorithm B3 with binary splitting and $x = 2^{33}$ (taking 205 hours on a dual-processor workstation "Nagisa").

Verification used Algorithm B1 with $x = 2^{34}$ (taking 269 hours).

The time to compute ln(2) (about 40 = 16 + 24 hours) is not included.

(ロ) (同) (E) (E) (E) (C)

Alexander Yee seems to hold the world record for the computation of γ (though records do not last long, so this may soon be obsolete). In March 2009 he computed the first 29 844 489 545 $\approx 2^{36}/\ln(10)$ decimal digits.

The computation was performed using Algorithm B3 with binary splitting and $x = 2^{33}$ (taking 205 hours on a dual-processor workstation "Nagisa").

Verification used Algorithm B1 with $x = 2^{34}$ (taking 269 hours).

The time to compute ln(2) (about 40 = 16 + 24 hours) is not included.

Note that B1 gives $34 \ln(2) + \gamma$ and B3 gives $33 \ln(2) + \gamma$, so it is possible to deduce both $\ln(2)$ and γ , but if we do this then we don't get an independent confirmation of the γ value.

◆□▶ ◆□▶ ◆三▶ ◆三▶ ◆□▶

Alexander Yee seems to hold the world record for the computation of γ (though records do not last long, so this may soon be obsolete). In March 2009 he computed the first 29 844 489 545 $\approx 2^{36}/\ln(10)$ decimal digits.

The computation was performed using Algorithm B3 with binary splitting and $x = 2^{33}$ (taking 205 hours on a dual-processor workstation "Nagisa").

Verification used Algorithm B1 with $x = 2^{34}$ (taking 269 hours).

The time to compute ln(2) (about 40 = 16 + 24 hours) is not included.

Note that B1 gives $34 \ln(2) + \gamma$ and B3 gives $33 \ln(2) + \gamma$, so it is possible to deduce both $\ln(2)$ and γ , but if we do this then we don't get an independent confirmation of the γ value.

For more details, history, and other constants, see http://
numberworld.org/nagisa_runs/computations.html.

Who was McMillan?

I met Ed McMillan when I was on sabbatical leave in Berkeley in 1977/8. At that time he had recently retired from Lawrence Berkeley Laboratory but still had an office there. He had seen my (first) paper on Euler's constant in *Math. Comp. 31 (1977)* and wanted to talk to me about possible improvements. Thus, I walked up the hill to LBL to talk to him, and our collaboration started.

э

Who was McMillan?

I met Ed McMillan when I was on sabbatical leave in Berkeley in 1977/8. At that time he had recently retired from Lawrence Berkeley Laboratory but still had an office there. He had seen my (first) paper on Euler's constant in *Math. Comp. 31 (1977)* and wanted to talk to me about possible improvements. Thus, I walked up the hill to LBL to talk to him, and our collaboration started.

Luckily I knew nothing about him at the time, or I might have been intimidated. He seemed to be just a scholarly old gentleman who was interested in Bessel functions.

< □ > < □ > < □ > □ =

To Sugard hiller to Confirting Sules Courtest Cini Mr. S. Dutte The most recent compartation of Sular contact & much a veries represent for the exponential integral, which was set egned to an sognifictio spession for the name $\int_{m}^{\infty} \frac{e^{-Y}}{2} J_{q} + \sum_{i=1}^{m} \frac{(-1)^{k}}{k!} \frac{m_{i}^{k}}{k!} = J_{m}m - Y \approx \frac{2^{2m}}{m} \left(1 - \frac{1}{m} + \frac{2!}{m!} - \frac{3!}{m!} - \dots \right) \quad ($ integral : The sequestion were in only print - so court ; if the new is tominded just before the mullest ten, the sure is but then the smallest term. In this case the conclust two is m! / " a Vite e", and the realty unsidently in the sale of the as too the end Therefor , if me desires I do a desired place , m about be chosen to be a Vi d to 10. The Also wire conveyer, but to match the presision of the compatible expression it meads the de animum to 15 4.32 in terms . It is an attempting series with a largest

13 9.3. Just, K.M. Conf. 21, 191-979 (1999)

1 × 4 = × 4 = × 1

·문▶·★ 문▶· · 문

The smallet them is at m = 2x+1, and has the value as 11/7 = - 1x. The renting inestantly in & ~ ~ 2 VT/x - +** Therefore, to get & to a driven flacer, X should be shown to be as 1/4 d har 10, i.d. half as lage to The walker of a meed in The exponential integral method. Both methods wed the some muchar of places (1/2 d) and the course survey of term (1/2 d he 10) in the contention of the agalitatic ancie. Both methods require waterity of a * (or a ") in the fatore multiplying the asymptotic series to Ved, and the Bund function method sequines in collition The (ad Vite where whe is a project spice) to that 3d/4 In out to match the pression stowed by the modely in A. (2), the win to Is (2) and So (2) must be more to to terms, where : 날 소 방 - 방 ~ . with the solution to a 2.31x. Thurks there resis must be surver to a 0.62 & 500

1 🕨 🖉 🖻 🕨 🖉 🕨 👘

3

time, the defined process. a configure of the two instances : defined more more of two states in separated integers 1.5 d 2.16 d (alo 3.14 d hall) Bund forther d 2.51 d da's 0.11 d hall

This great advantage is aroundant combinistic by the fact that the Bend forten method require the amounting of the two avies Io(x) and So(x), but these have in common the freta (1/1) 2 / (del) 2 , to the summeties can be carried on in handled. I for The factor (1+1 +3 + ... +) can be altered, with with (an additional recipiered address of inche sty. Job 8() # with (but a get and it - are 77100/1. - the) with the construction of the state Results with a heart combate.

(国) ▲ 国) ト 国 (の) Q ()

Later I discovered that McMillan had worked at Los Alamos during the second world war, and was famous for discovering the first trans-uranium element (neptunium) in 1940, soon followed by the discovery of plutonium with Seaborg.

伺 ト イ ヨ ト イ ヨ ト

크

Later I discovered that McMillan had worked at Los Alamos during the second world war, and was famous for discovering the first trans-uranium element (neptunium) in 1940, soon followed by the discovery of plutonium with Seaborg.

He shared the 1951 Nobel prize in Chemistry with Seaborg for this work.

□→★□→★□→ □

Later I discovered that McMillan had worked at Los Alamos during the second world war, and was famous for discovering the first trans-uranium element (neptunium) in 1940, soon followed by the discovery of plutonium with Seaborg.

He shared the 1951 Nobel prize in Chemistry with Seaborg for this work.

In 1945 he discovered (independently of Veksler) the theory of the synchrotron, making possible the construction of modern particle accelerators.

日本・モト・モト・モ

Later I discovered that McMillan had worked at Los Alamos during the second world war, and was famous for discovering the first trans-uranium element (neptunium) in 1940, soon followed by the discovery of plutonium with Seaborg.

He shared the 1951 Nobel prize in Chemistry with Seaborg for this work.

In 1945 he discovered (independently of Veksler) the theory of the synchrotron, making possible the construction of modern particle accelerators. Modified Bessel functions of fractional order occur in this theory.

御 と く き と く き と … き

Later I discovered that McMillan had worked at Los Alamos during the second world war, and was famous for discovering the first trans-uranium element (neptunium) in 1940, soon followed by the discovery of plutonium with Seaborg.

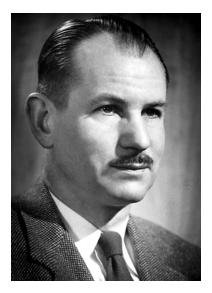
He shared the 1951 Nobel prize in Chemistry with Seaborg for this work.

In 1945 he discovered (independently of Veksler) the theory of the synchrotron, making possible the construction of modern particle accelerators. Modified Bessel functions of fractional order occur in this theory.

According to the obituary by Jackson and Panofsky, McMillan's last published paper was the one that he wrote with me on the computation of Euler's constant. It was published in *Math. Comp.* 34 (1980).

▲□ ▶ ▲ 臣 ▶ ▲ 臣 ▶ ● 臣 ● の Q @

Edwin Mattison McMillan 1907–1991 (about 1950)



< □ > < □ > < □ > .

크

Another photo (much as I remember him in 1977)



Edin An Horalla

Richard Brent Ramanujan and Euler's Constant

문 > · · 문 > · · 문

Prehistory - Riemann 1855

III. Zur Theorie der Nobili'schen Farbenringe 59 $\int_{\frac{1}{\sqrt{tt-1}}}^{\infty} \frac{e^{-2qt}dt}{\sqrt{tt-1}} \operatorname{durch} f(q), \int_{\frac{1}{\sqrt{1-tt}}}^{1} \frac{e^{-2qt}dt}{\sqrt{1-tt}} \operatorname{durch} \varphi(q)$ reichnet $b_n = c_n f\left(n \frac{\pi}{AB} r\right) + \gamma_n \varphi\left(n \frac{\pi}{AB} r\right).$ Die Entwicklung nach steigenden Potenzen von q giebt $f(q) = \sum_{m \mid m} \frac{q^{2m}}{m! m!} (\Psi(m) - \log q)$ $\varphi(q) = \pi \sum \frac{q^{3m}}{m!m!};$ wird also f(q) für q = 0 unendlich und damit u'' für r = 0 station bleibe, muss $c_s = 0$ sein; γ_s ergiebt sich dann aus (4) gleich $-\frac{4 \sin n \frac{\pi}{2\beta} \alpha f' \left(n \frac{\pi}{4\beta} c\right)}{\beta \omega' \left(n \frac{\pi}{2\beta} c\right)},$ mithin $u = \mathcal{E}^{n} \sin n \frac{\pi}{2\beta} x \frac{4\sin n \frac{\pi}{2\beta} \alpha}{\beta} \left\{ f\left(n \frac{\pi}{4\beta} r\right) - \varphi\left(n \frac{\pi}{4\beta} r\right) \frac{f'\left(n \frac{\pi}{4\beta} c\right)}{\sigma'\left(n \frac{\pi}{\alpha} c\right)} \right\}$ über alle positiven ungeraden Werthe von n ausgedehnt. Zur Berechnung von f(q) und $\varphi(q)$ können für grosse Werthe von q die halbconvergenten Reihen $f(q) = e^{-2q} \sqrt{\frac{\pi}{4q}} \sum_{i=1}^{\infty} (-1)^m \frac{(1 \cdot 3 \dots 2m - 1)^i}{m! \ (16q)^m},$ $\varphi(q) = e^{2q} \sqrt{\frac{\pi}{4q}} \sum_{i=1}^{r} \frac{(1 \cdot 3 \dots 2m - 1)^{i}}{m! (16q)^{m}}$ benutzt werden, welche indess ihren Werth nur bis auf Bruchtheile von der Ordnung der Grösse e-49 geben; genügt diese Genauigkeit nicht, so ist es wohl am zweckmässigsten die Entwicklungen nach steigenden Potenzen von g anzuwenden. Für hinreichend grosse Werthe von $\frac{r}{s}$ erhält man also mit Vernachlässigung von Grössen von der Ordnung der Grösse $e^{-b\frac{\pi}{2\beta'}}$

- 同下 - ヨト - ヨト

э

Conclusion

We did not succeed in proving that γ is irrational, but the quest was worthwhile because it provided the motivation:

- to read Ramanujan's papers and Notebooks;
- to meet and collaborate with Ed McMillan;
- to learn more about Bessel functions (a topic of classical mathematics that should be better known).

< □→ < □→ < □→ □ □

Conclusion

We did not succeed in proving that γ is irrational, but the quest was worthwhile because it provided the motivation:

- to read Ramanujan's papers and Notebooks;
- to meet and collaborate with Ed McMillan;
- to learn more about Bessel functions (a topic of classical mathematics that should be better known).

Finally, here is a nice integral for $K_0(x)$, x > 0:

$$\mathcal{K}_0(x) = \int_0^\infty \frac{\cos(xt)}{\sqrt{1+t^2}} \,\mathrm{d}t.$$

< □→ < □→ < □→ - □

Conclusion

We did not succeed in proving that γ is irrational, but the quest was worthwhile because it provided the motivation:

- to read Ramanujan's papers and Notebooks;
- to meet and collaborate with Ed McMillan;
- to learn more about Bessel functions (a topic of classical mathematics that should be better known).

Finally, here is a nice integral for $K_0(x)$, x > 0:

$$\mathcal{K}_0(x) = \int_0^\infty \frac{\cos(xt)}{\sqrt{1+t^2}} \,\mathrm{d}t.$$

We could use this with numerical quadrature to compute γ , but it would be unlikely to give a fast algorithm.

▲御▶ ▲理▶ ▲理▶ 二理

References

1. Bruce C. Berndt, *Ramanujan's Notebooks,* Parts I–III, Springer-Verlag, New York, 1985–1991.

2. Richard P. Brent and Edwin M. McMillan, Some new algorithms for high-precision computation of Euler's constant, *Mathematics of Computation* **34** (1980), 305–312.

3. Richard P. Brent, An asymptotic expansion inspired by Ramanujan, *Australian Mathematical Society Gazette* 20 (December 1993), 149–155. arXiv:1004.5506v1.

4. Richard P. Brent, Ramanujan and Euler's constant, *Proceedings of Symposia in Applied Mathematics*, Vol. 48 AMS, Providence, Rhode Island, 1994, 541–545.

(本間) (本語) (本語) (語)

References cont.

5. J. D. Jackson and W. K. H. Panofsky, Edwin Mattison McMillan 1907-1991, *Biographical Memoirs of the National Academy of Sciences (USA)*, 69 (1996), 214–241.

6. Thom Mulders, On short multiplications and divisions, *Applicable Algebra in Engineering, Communication and Computing*, 11, 1 (2000), 69–88.

7. Frank W. J. Olver, *Asymptotics and Special Functions*, Academic Press, New York, 1974.

8. Srinivasa Ramanujan, A series for Euler's constant γ , *Messenger of Mathematics* 46 (1917), 73–80.

9. G. F. B. Riemann, Zur Theorie der Nobili'schen Farbenringe, *Poggendorff's Annalen der Physik und Chemie*, 95 (1855), 130–139. Reprinted in *Bernhard Riemann's Gesammelte Mathematische Werke*, Dover, NY, 1953, 55–62.

□ > < E > < E > □ E